

# Investigating Coating Materials and Processes for FUV to NIR

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## Astrophysics

Excerpts from a report from Prof. Paul Scowen, ASU

- It has been recognized that at mid to far ultraviolet wavelengths (90 <  $\lambda$  < 300 nm), it is possible to detect and measure important astrophysical processes, which can shed light into the physical conditions of many environments of interest.
- For example, in the local interstellar medium (LISM) all but two (Ca II H and K lines) of the key diagnostic of resonance lines are in the ultraviolet (Redfield 2006).
- In addition to the fruitful science areas that ultraviolet spectroscopy has contributed since the early 1970s, France et al. (2013) have emphasized the role of ultraviolet photons in the photodissociation and photochemistry of H2O and CO2 in terrestrial planet atmospheres, which can influence their atmospheric chemistry, and subsequently the habitability of Earth-like planets.
- Similarly, new areas of scientific interest are the detection and characterization of the hot gas between galaxies and the role of the intergalactic medium (IGM) in galaxy evolution (Shull et al., 2012).

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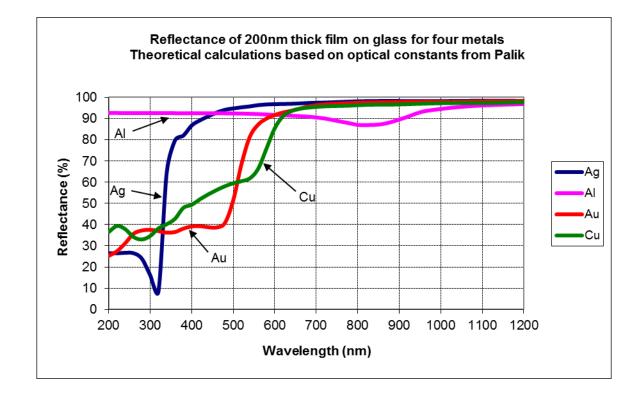
#### **COR Program Goals**

- The NASA Cosmic Origins Program Annual Technology Report (COR Technology Needs, Table 7, Item 8.1.3., page 43, Oct 2011) defined the primary goal that we have adopted for this project: "Development of UV coatings with high reflectivity (>90-95%), high uniformity (<1-0.1%), and wide bandpasses (~100 nm to 300-1000 nm)".
- High reflectivity coatings covering the 100-120nm spectral region is considered important for studying intergalactic matter (IGM). The COPAG assessed the degree of difficulty to achieving this as very high.
- Void-free thin films of absorption-free materials are required to protect and maintain high reflectivity and durability of aluminum mirrors in laboratory and prelaunch environments. Precisely controllable and scalable deposition process is also required to produce such coatings on large telescope mirrors.

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#### Background



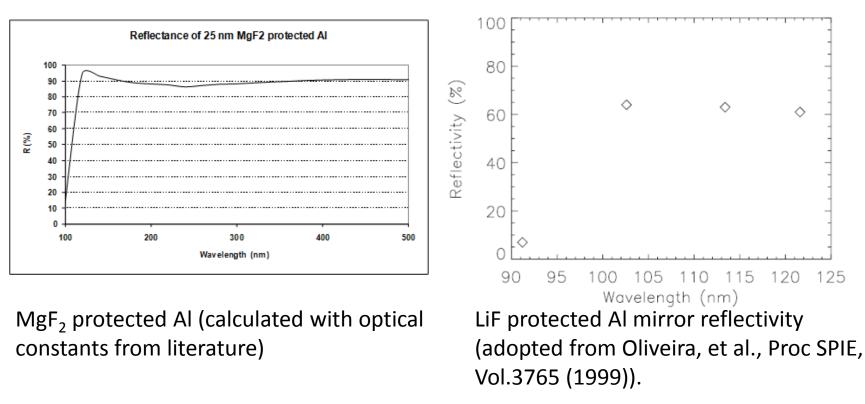
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#### Background

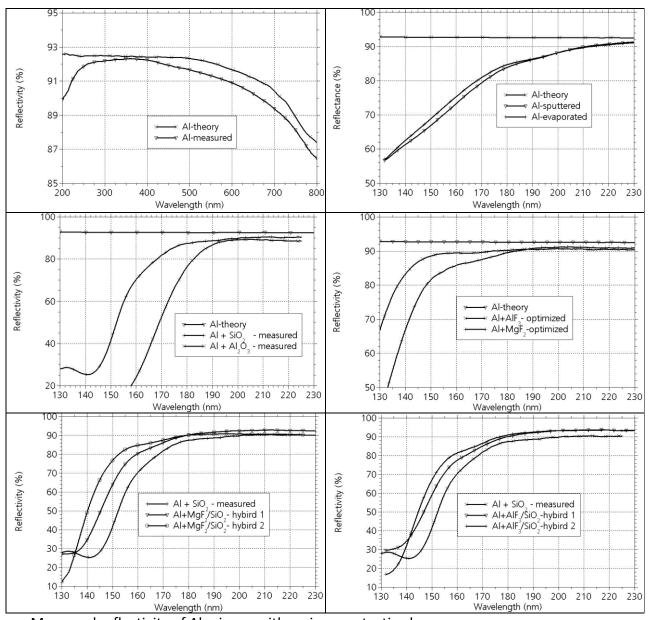
Hubble Telescope MgF<sub>2</sub> coated Al mirror >115nm through visible wavelengths

#### FUSE (Far Ultraviolet Spectroscopic Explorer)



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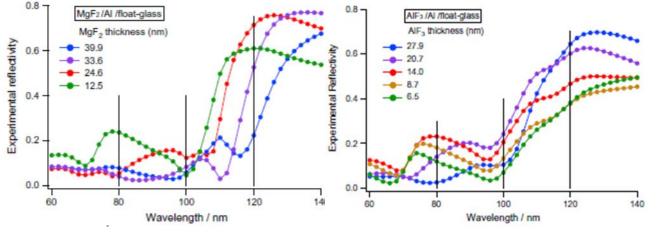




Measured reflectivity of Al mirrors with various protective layers Figures adopted from Yang, Gatto and Kaiser, Proc SPIE vol.5693, (2005)



## $MgF_2$ and $AIF_3$ on AI



 $MgF_2/Al$  coating reflectivity in the DUV [Bridou, et al (2010)]

 $AIF_3/AI$  coating reflectivity in the DUV [Bridou, et al (2010)]

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#### **Candidate Materials**

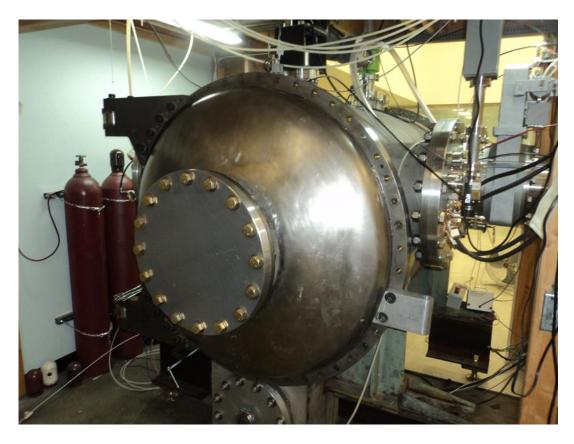
Several fluorides: CaF<sub>2</sub>, LiF, MgF<sub>2</sub>, LaF<sub>3</sub>, AlF<sub>3</sub>, Na<sub>3</sub>AlF<sub>6</sub>, YbF<sub>3</sub> and GdF<sub>3</sub>

- Produced single layer coatings of MgF<sub>2</sub>, LiF, AlF<sub>3</sub>, LaF<sub>3</sub>, Na<sub>3</sub>AlF<sub>6</sub> and GdF<sub>3</sub> with conventional thermal evaporation at pressures in the range of 5x10<sup>-7</sup> to 1x10<sup>-6</sup> Torr and temperatures in the range of 180 to 200C.
- Coatings were prepared on fused silica and silicon substrates. UV grade fused silica substrates of 2 inch diameter were specially prepared with both sides polished and one side polished regions over which coatings could be prepared in the same run to enable multiple experiments / measurements with one sample.
- Detailed characterization of these coatings is in progress.

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#### 1.2 m Coating Chamber



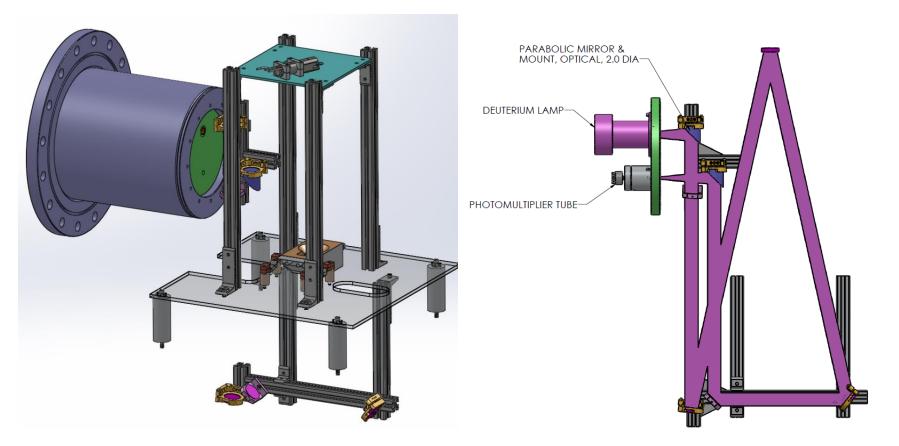
# Coating chamber employed to produce various coatings for initial experiments with conventional coating techniques

Courtesy: Zecoat Corp.

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## Reflectometer (work in progress)

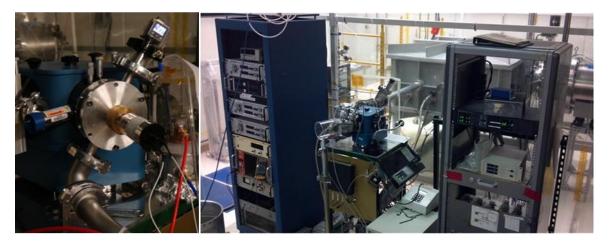


#### Courtesy: Zecoat Corp.

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#### FUV measurements at Univ. of Colorado

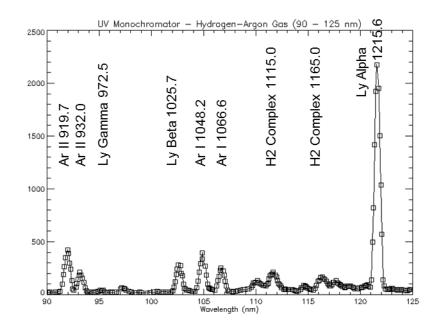


FUV measurement system at Univ. of Colo. (Courtesy: Prof. James Green). The light sources include sealed Pt-Ne and deuterium discharge lamps, and windowless gas discharge systems. These can provide a host of emission lines from 400 – 2000 Å, and continuum emission >1608 Å. The monochromator is a normal incidence McPherson - it's selectable bandwidth can be scanned over the full 400 – 2000 Å range. The vacuum tank is operated in a class 2000 clean tent.

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#### FUV measurements at Univ. of Colorado

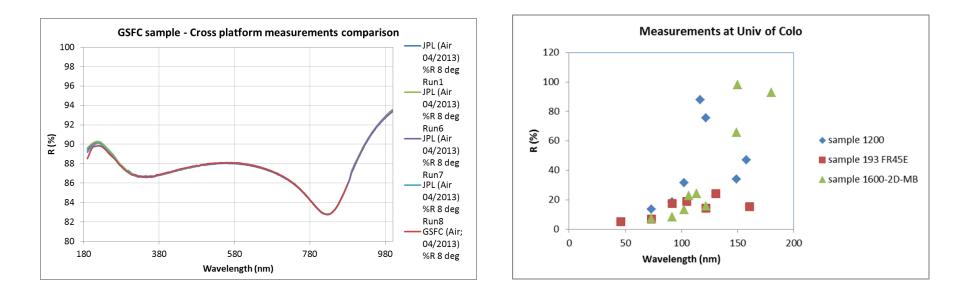


# Representative measurement results from Univ. of Colorado FUV measurement system indicating its current functionality for further work.

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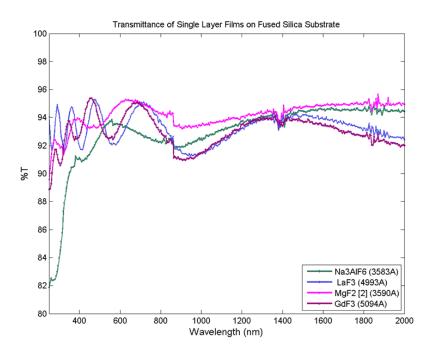
#### **Commercial samples measured**



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#### Single layer coating samples Measured Transmittance



Transmittance of Single Layer Films on Fused Silica Substrate 100 98 %T 88 86 84 LiF [2] (3964A) LiF - [Ar+] (4678A) 82 MgF2 [2] (3590A) JF3 [2] (4700A) 80 400 600 800 1000 1200 1400 2000 1600 1800 Wavelength (nm)

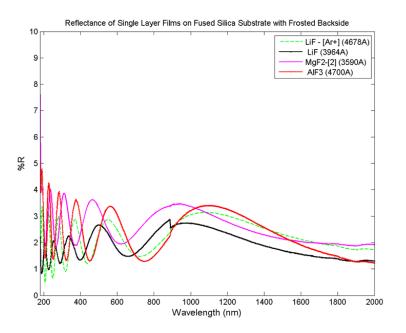
Transmittance spectra of single layer coatings of  $GdF_3$ ,  $MgF_2$ ,  $LaF_3$ , and  $Na_3AIF_6$  on uv grade fused silica substrate. The numbers in parenthesis indicate thickness in Angstroms. Angle of incidence: 0 deg.

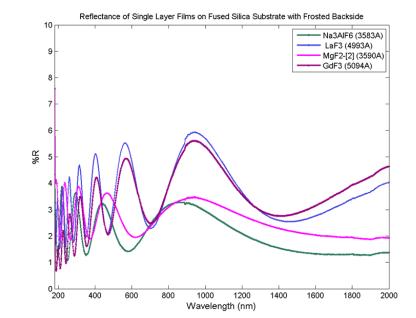
Transmittance spectra of single layer coatings of  $AIF_3$ ,  $MgF_2$ , and LiF on uv grade fused silica substrate. The numbers in parenthesis indicate thickness in Angstroms. Angle of incidence: 0 deg.

Measurements done with a Perkin Elmer UV-VIS Spectrophotometer



#### Single layer coating samples Measured Reflectance



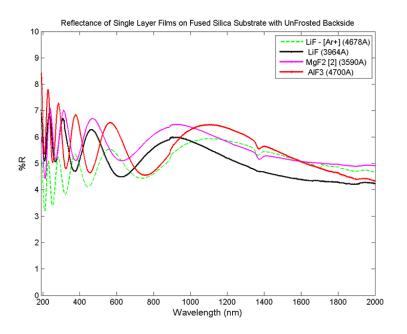


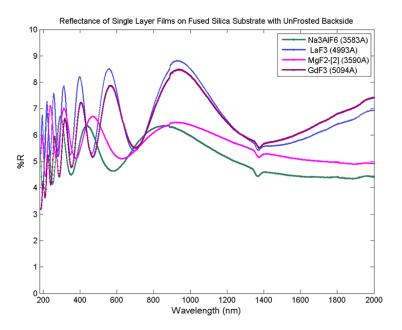
Reflectance spectra of single layer coatings of  $MgF_2$ , LiF, and  $AlF_3$  on uv grade fused silica substrate with frosted backside. The numbers in parenthesis indicate thickness in Angstroms. Angle of incidence: 8 deg Reflectance spectra of single layer coatings of  $GdF_3$ , MgF<sub>2</sub>, LaF<sub>3</sub>, and Na<sub>3</sub>AlF<sub>6</sub> on uv grade fused silica substrate with frosted backside. The numbers in parenthesis indicate thickness in Angstroms. Angle of incidence: 8 deg

Measurements done with a Perkin Elmer UV-VIS Spectrophotometer with a Universal Reflectance Accessory



#### Single layer coating samples Measured Reflectance



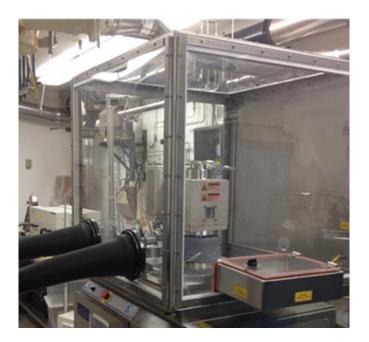


Reflectance spectra of single layer coatings of  $AIF_3$ ,  $MgF_2$ , and LiF on uv grade fused silica substrate with unfrosted backside. The numbers in parenthesis indicate thickness in Angstroms. Angle of incidence: 8 deg Reflectance spectra of single layer coatings of  $LaF_3$ ,  $GdF_3$ ,  $MgF_2$  and  $Na_3AIF_6$  on uv grade fused silica substrate with unfrosted backside. The numbers in parenthesis indicate thickness in Angstroms. Angle of incidence: 8 deg

Measurements done with a Perkin Elmer UV-VIS Spectrophotometer with a Universal Reflectance Accessory



#### ALD system at JPL

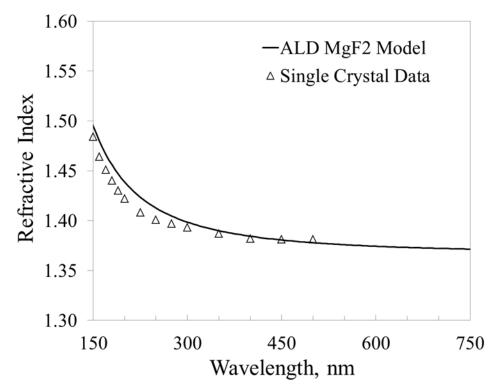


#### ALD coating system at JPL

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## MgF<sub>2</sub> Film by ALD



Refractive index of MgF<sub>2</sub> film recently grown with ALD process at JPL. MgF<sub>2</sub> deposited with Mg(CpEt)2 and HF at 200 °C, ~100 nm; Model is ellipsometric best fit for measurement over the range 1.5 - 6.5 eV; Film appears transparent in this range, apparent k=0; Single crystal data from (Williams, Applied Optics 18, 1979)



## Plans for the coming year (FY14)

- Produce protected Al mirror samples with chosen protective layers
- Develop and perform environmental tests (humidity and thermal cycling) to establish protection of Al and its reflectivity, particularly in the deep UV.
  Measure R, before and after environmental tests, and characterize the surface microscopically.
- Develop and optimize ALD process for absorption-free thin MgF<sub>2</sub> coatings, and MgF<sub>2</sub> protected Al mirrors.
- Compare the performance characteristics of protective coatings made by conventional techniques (e.g., IAD) and by ALD



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Team Members: Stuart Shaklan, JPL Nasrat Raouf, JPL Shouleh Nikzad, JPL John Hennessy, JPL Michael Ayala, JPL

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